

# Astronomy 1 – Introductory Astronomy

## Spring 2014

Equations and constants and conversion factors for the final exam

$$1 \text{ kilometer (km)} = 1000 \text{ meters (m)}$$

$$1 \text{ year (yr)} = 3.15 \times 10^7 \text{ seconds (s)}$$

$$1 \text{ day} = 86,400 \text{ s}$$

$$1 \text{ Watt (W)} \text{ is } 1 \text{ Joule (J)} \text{ per second (} W = J/s)$$

$$1 \text{ electron Volt (eV)} = 1.60 \times 10^{-19} \text{ J}$$

1 Hertz (Hz) is 1 cycle/second (1/s), the fundamental unit of frequency

$$1 \text{ Newton (unit of force)} = \text{kg} \cdot \text{m}/\text{s}^2$$

$$1 \text{ Joule (unit of energy)} = \text{kg} \cdot \text{m}^2/\text{s}^2$$

$$c = 3.0 \times 10^8 \text{ m/s}$$

$$h = 6.63 \times 10^{-34} \text{ Js}$$

$$k = 1.38 \times 10^{-23} \text{ J/K}$$

$$G = 6.67 \times 10^{-11} \text{ Nm}^2/\text{kg}^2$$

$$\sigma = 5.67 \times 10^{-8} \text{ W/m}^2/\text{K}^4$$

$$m_p = 1.67 \times 10^{-27} \text{ kg}$$

$$H_o = 22 \text{ km/s/Mly}$$

$$1 \text{ parsec (pc)} = 3.09 \times 10^{16} \text{ m} = 3.26 \text{ light years}$$

$$1 \text{ A.U.} = 1.5 \times 10^{11} \text{ m}$$

$$R_{\text{sun}} = 7.0 \times 10^8 \text{ m}$$

$$R_{\text{Jupiter}} = 7.1 \times 10^7 \text{ m}$$

$$R_{\text{earth}} = 6.4 \times 10^6 \text{ m}$$

$$R_{\text{moon}} = 1.7 \times 10^6 \text{ m}$$

$$L_{\text{sun}} = 3.8 \times 10^{26} \text{ W}$$

$$M_{\text{sun}} = 2.0 \times 10^{30} \text{ kg}$$

$$M_{\text{Jupiter}} = 1.9 \times 10^{27} \text{ kg}$$

$$M_{\text{earth}} = 5.97 \times 10^{24} \text{ kg}$$

$$\theta = \frac{57.3L}{d}$$

$$v = \frac{d}{t}$$

$$B = \frac{L}{4\pi d^2}$$

$$E = hf$$

$$\lambda f = c$$

$$\lambda_{max} = \frac{2.9 \times 10^6 \text{ nmK}}{T}$$

$$B_{surf} = \sigma T^4$$

$$\frac{v_r}{c} = \frac{\lambda_{shift} - \lambda_{rest}}{\lambda_{rest}}$$

$$F = ma$$

$$F_g = \frac{GMm}{r^2}$$

$$a_{cent} = \frac{v^2}{r}$$

$$v^2 = \frac{GM}{r}$$

$$p^2 = \frac{4\pi^2 a^3}{G(M + m)}$$

$$p^2 = \frac{a^3}{M}$$

$$M_1 r_1 = m_2 r_2$$

$$P = nkT$$

$$\rho = mn$$

$$p = \frac{1}{d}$$

$$L = 4\pi R^2 \sigma T^4$$

$$E = mc^2$$

$$A_{sphere} = 4\pi R^2$$

$$V_{sphere} = \frac{4}{3}\pi R^3$$

$$C_{circle} = 2\pi R$$

$$PE = mgh$$

$$KE = \frac{1}{2}mv^2$$

$$\text{fraction of light blocked} = (R_p/R_s)^2$$

$$M_1 v_1 = m_2 v_2$$

$$v_p = \frac{2\pi a_p}{p_p}$$

$$M_p = \frac{M_s v_s P_p}{2\pi a_p}$$

density = mass/volume

signal = brightness\*exposure time\*collecting area\*sensitivity

telescope collecting area =  $\pi(D/2)^2$

$$\frac{\text{current amount}}{\text{original amount}} = \left(\frac{1}{2}\right)^{t/t_{half}}$$

$$M_{enclosed} = \frac{rv^2}{G}$$

$$v = H_0 d$$

$$\text{age} = 1/H_0$$

$$\text{redshift, } z = \frac{v}{c} = \frac{\lambda_{observed} - \lambda_{rest}}{\lambda_{rest}}$$

$$1 + z = \frac{\lambda_{observed}}{\lambda_{rest}}$$